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Designed By James Hultquist-Todd 4 Weights Released in 2020

Reflating Fanlight Tombac Cracton,

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Hugoslav Bebrocks Aborably Skreighs

Signate Supawn An-Built Almoney

Boubles Ambush Cushaw Bincheb

CHYORE KODI

Bundityy Undergrab Filter-Paper Bury the hatchet Accidents will happen Ribble wrapped up in an enigma

Men's evil manners live in brass; their virtues we write in water

Clustery Extrusion 到ot-妈anger Selection 2373 A fly in the ointment Fill go to the foot of our stairs

Th, they have slain the Earl of Moray and Lady Mondegreen

Wisthuay Rehabilitatie Bubble and squeak Parting shot / Parthian shot Eye of newt and toe of frog, wool of bat and tongue of dog

nhues nfta dia Imperishable at the at Norwegian: 800-357-4159

A journey of a thousand miles begins with a single step

UPPERCASE

abcbefghijklmnopqrstubwxyz

LOWERCASE

0123456789 0123456789 \$¢±¥€∮f

NUMBERS AND CURRENCY

Mathematical Symbols

Punctuation

#%%%%®®®®%~

Symbols

There is an occasional star, like chi Carinae, whose spectrum consists almost wholly of bright lines, in general bearing no apparent relationship to the bright lines in the spectra of the gaseous nebulae except that the hydrogen lines are there, as they are almost everywhere. There is reason to believe that such a spectrum indicates the existence of a very extensive and very hot atmosphere surrounding the main body, or core, of the star in question. This particular star is remarkable in that it has undergone great changes in brilliancy and is located upon a background of nebulosity. The chances are strong that the star has rushed through the nebulosity with high rate of speed and that the resulting bombardment of the star has expanded and intensely heated its atmosphere.

There are the Wolf-Rayet stars, named from the French astronomers who discovered the first three of this class, whose spectra show a great variety of combinations of continuous spectrum and bright bands. We believe that the continuous spectrum in such a star comes from the more condensed central part, or core, and that the bright-line light proceeds from a hot atmosphere extending far

out from the core.

The great majority of the stars have spectra which are continuous, except for the presence of bark or absorption lines: a few lines in the very blue stars, and an